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Origin of the volcanic-hosted Yamansu Fe deposit, Eastern Tianshan, NW China: constraints from pyrite Re-Os isotopes, stable isotopes, and in situ magnetite trace elements

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Received: 14 June 2017 /Accepted: 15 January 2018 /Published online: 31 January 2018 \odot Springer-Verlag GmbH Germany, part of Springer Nature 2018

Abstract

The Yamansu Fe deposit (32 Mt at 51% Fe) in the Eastern Tianshan Orogenic Belt of NW China is hosted in early Carboniferous volcano-sedimentary rocks and spatially associated with skarn. The paragenetic sequence includes garnetdiopside (I), magnetite (II), hydrous silicate-sulfide (III), and calcite-quartz (IV) stages. Pyrite associated with magnetite has a Re-Os isochron age of 322 ± 7 Ma, which represents the timing of pyrite and, by inference, magnetite mineralization. Pyrite has $\delta^{34}S_{\text{VCDT}}$ values of −2.2 to +2.9‰, yielding $\delta^{34}S_{\text{H2S}}$ values of −3.1 to 2‰, indicating the derivation of sulfur from a magmatic source. Calcite from stages II and IV has $\delta^{13}C_{\text{VPDB}}$ values from − 2.5 to − 1.2‰, and − 1.1 to 1.1‰, and $\delta^{18}O_{VSMOW}$ values from 11.8 to 12.0‰ and −7.7 to −5.2‰, respectively. Calculated $\delta^{13}C$ values of fluid $CO₂$ and water $\delta^{18}O$ values indicate that stage II hydrothermal fluids were derived from magmatic rocks and that meteoric water mixed with the hydrothermal fluids in stage IV. Some ores contain magnetite with obvious chemical zoning composed of dark and light domains in BSE images. Dark domains have higher Mg, Al, Ca, Mn, and Ti but lower Fe and Cr contents than light domains. The chemical zoning resulted from a fluctuating fluid composition and/or physicochemical conditions (oscillatory zoning), or dissolution-precipitation (irregular zoning) via infiltration of magmatic-hydrothermal fluids diluted by late meteoric water. Iron was mainly derived from fluids similar to that in skarn deposits.

Keywords Re-Os geochronology . Trace elements . Magnetite . Dissolution-reprecipitation . Yamansu Fe deposit

Editorial handling: R. Hu

Electronic supplementary material The online version of this article (<https://doi.org/10.1007/s00126-018-0794-4>) contains supplementary material, which is available to authorized users.

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Introduction

The Paleozoic Tianshan Orogenic Belt (TOB), NW China, has been a focus of mineral exploration in China (Wang et al. [2006a\)](#page-21-0). Particularly, the Chagangnuoer, Beizhan, Dunde, and Zhibo Fe deposits are well known in the western part of the TOB (Hou et al. [2014a;](#page-18-0) Zhang et al. [2014\)](#page-21-0) and the Yamansu, Heifengshan, Shuangfengshan, and Shaquanzi Fe deposits in its eastern part (Gao et al. [1993](#page-18-0); Che et al. [1994;](#page-18-0) Qin et al. [2002;](#page-20-0) Wang et al. [2006a](#page-21-0); Huang et al. [2013a](#page-19-0), [2014a,](#page-19-0) [b\)](#page-19-0). These deposits were thought to be submarine volcanogenic in origin because they are hosted in volcanic rocks (Jiang [1983;](#page-19-0) Jiang and Wang [2005](#page-19-0)). They are also considered to be related to skarnization because of extensive skarn alteration and associated with concealed intrusions (Mao et al. [2005;](#page-19-0) Pirajno [2010\)](#page-20-0), although they are different from typical skarn deposits in that they are not directly in contact with intrusions (Einaudi et al. [1981](#page-18-0)). The Yamansu deposit contains 32 Mt at

51% Fe (Mao et al. [2005\)](#page-19-0) and is one of the largest Fe deposits in the eastern TOB. This deposit has many geological, mineralogical, and geochemical features similar to other alkali-calcic, hydrothermally altered volcanic rock-hosted Fe deposits, which are the subject of a long-standing debate concerning a possible magmatic or hydrothermal origin, or a combination of both (Henriquez and Martin [1978](#page-18-0); Nyström and Henríquez [1994;](#page-20-0) Rhodes et al. [1999;](#page-20-0) Sillitoe and Burrows [2002;](#page-20-0) Nyström et al. [2008](#page-20-0); Dare et al. [2015;](#page-18-0) Zhang et al. [2015;](#page-21-0) Tornos et al. [2016](#page-21-0); Günther et al. [2017](#page-18-0); Tang et al. [2017;](#page-20-0) Zheng et al. [2017\)](#page-21-0). For example, recent studies of in situ trace elements and iron isotopes of magnetite from the Zhibo and Chagangnuoer Fe deposits demonstrated a combined orthomagmatic and skarn process for the formation of these deposits (Günther et al. [2017\)](#page-18-0). However, the timing of mineralization, origin and evolution of hydrothermal fluids, and iron enrichment mechanism of the Yamansu deposit are still lacking. Therefore, detailed studies of the Yamansu deposit will provide important constraints on the origin of volcanic rockhosted Fe deposits in the TOB.

In this study, we describe the field relationships, paragenetic sequence, and mineralization styles of the Yamansu deposit. Pyrite Re-Os geochronology is used to date pyrite that is associated with magnetite. Carbon and oxygen isotope compositions of calcite and sulfur isotope composition of pyrite, together with in situ trace element composition of magnetite, are used to constrain the origin and evolution of the hydrothermal fluids and the mechanism of iron enrichment in the Yamansu deposit. The implications for the regional Fe(-Cu) metallogeny in the ETOB are also discussed.

Regional geology

The Central Asian Orogenic Belt (CAOB) is the largest Phanerozoic orogen in the world, extending \sim 7000 km from west to east, and is a complex collage of continental fragments, island arc assemblages, remnants of oceanic crust, and continental margins between the Siberian Craton to the north and the Tarim-North China Craton to the south (Fig. [1](#page-2-0)a). It was formed by multiple subduction-accretion and collision processes from the Neoproterozoic to the Late Paleozoic (Sengör and Burtman [1993;](#page-20-0) Jahn et al. [2000;](#page-19-0) Windley et al. [2007;](#page-21-0) Xiao et al. [2008](#page-21-0)). The TOB in the southern CAOB extends from northern Xinjiang to Gansu for more than 3000 km (Fig. [1b](#page-2-0)). The eastern part of the TOB (ETOB) comprises four main tectonic units from north to south, the Dananhu-Tousuquan, Kanggurtag, Aqishan-Yamansu, and Central Tianshan belts (Fig. [1c](#page-2-0)) (He et al. [1994;](#page-18-0) Gao et al. [1998;](#page-18-0) Qin et al. [2002](#page-20-0); Xiao et al. [2004](#page-21-0); Wang et al. [2006b\)](#page-21-0).

The Dananhu-Tousuquan belt is composed of Devonian to Carboniferous island arc volcanic rocks. Devonian rocks, basaltic to andesitic volcanic rocks, are locally associated with carbonate and calcareous mudstone. Overlying these rocks is the late Carboniferous Tuwu Formation, composed of greywacke and tuff intercalated with carbonates. The Tuwu Formation is overlain by Permian basalt, tuff and volcanic breccia, and Jurassic terrestrial clastic rocks (Mao et al. [2005\)](#page-19-0).

In the Kanggurtag belt, there are mylonite, tectonic lenses, and breccias formed along early compressional faults (Qin et al. [2002;](#page-20-0) Xu et al. [2003](#page-21-0)). The Carboniferous strata consist of fine-grained sandstone and carbonaceous argillite. The Kanggurtag belt contains ophiolitic fragments and was thought to mark a suture zone separating the Siberia craton to the north from the Tarim block to the south (Ji et al. [1994;](#page-19-0) Yang et al. [1996](#page-21-0); Ma et al. [1997](#page-19-0); Zhou et al. [2001](#page-21-0)). The Kanggurtag belt was also thought to be an inter-arc basin between the Dananhu-Tousuquan and Yamansu arcs (Ma et al. [1997](#page-19-0); Shu et al. [2002](#page-20-0)).

The Aqishan-Yamansu belt is separated by the Kushui fault from the Kanggurtag belt to the north and by the Shaquanzi fault from Central Tianshan belt to the south. In the Aqishan-Yamansu belt, the lower Carboniferous Yamansu Formation is composed of a 5-km-thick succession of bimodal volcanic rocks, the middle Carboniferous Shaquanzi Formation of the clastic rocks and andesitic tuff, and the upper Carboniferous Tugutublak Formation of intercalated carbonate and volcanic rocks. There are also Permian marine and terrestrial clastic rocks intercalated with bimodal volcanic rocks and carbonates (Mao et al. [2005](#page-19-0)). Dacite and rhyolite of the Yamansu Formation in the east, middle, and western parts of ETOB have zircon U-Pb ages of 348.0 ± 1.7 , 335.9 ± 2.4 , and 334.0 ± 2.5 Ma, respectively (Luo et al. [2012](#page-19-0)). Rhyolite of the Shaquanzi Formation has a zircon U-Pb age of $320 \pm$ 1.2 Ma (Li et al. [2011\)](#page-19-0). This belt was thought to be an island arc belt (Yan [1985](#page-21-0); Ma et al. [1997;](#page-19-0) Yang et al. [1999](#page-21-0)). However, it has also been interpreted to be an early Carboniferous rift zone because of widespread intraplate riftrelated volcanic rocks (Xiao et al. [1992;](#page-21-0) Qin et al. [2002;](#page-20-0) Xia et al. [2003\)](#page-21-0). Iron deposits in this belt include the larger Yamansu deposit, the smaller Heifengshan, Shuangfengshan, Shaquanzi, Bailingshan, and Hongyuntan deposits (Fig. [1c](#page-2-0)).

The Central Tianshan belt in the ETOB has been regarded as a composite volcanic arc, composed of Precambrian basement rocks of amphibolite facies, overlain by calc-alkaline basaltic andesite, volcanoclastic rocks, intruded by minor Itype granite and granodiorite (Xiao et al. [2004](#page-21-0)). The Precambrian basement consists of gneiss, quartz schist, migmatite, and marble. Iron deposits in this belt are represented by the Precambrian metamorphic rocks hosted Tianhu deposit (Huang et al. [2015;](#page-19-0) Wu et al. [2016](#page-21-0)) and Permian alkaline gabbro complex hosted Weiya magmatic Fe-Ti-V deposit (Wang et al. [2005;](#page-21-0) Zhang et al. [2005\)](#page-21-0).

Large volumes of Carboniferous to Permian granitic rocks in the ETOB include quartz diorite, granodiorite, monzonite, biotite granite, and K-feldspar granite (Tong et al. [2010;](#page-20-0) Zhou

Fig. 1 a Location of the study area in the Central Asia Orogenic Belt (modified from Zhang et al. [2009\)](#page-21-0). b Sketch map showing geologic units of northern Xinjiang, NW China (modified from Xie et al. [2012](#page-21-0)). c Regional geological map of the Eastern Tianshan Orogenic Belt

(modified from Wang et al. [2006b\)](#page-21-0). Faults: ①—Kanggurtag-Huangshan fault, ②—Kushui fault, ③—Shaquanzi fault. Iron deposits: 1—Hongyuntan, 2—Tieling, 3—Bailingshan, 4—Yamansu, 5— Heifengshan, 6—Shuangfengshan, 7—Shaquanzi, 8—Weiya, 9—Tianhu

et al. [2010](#page-21-0)). There are numerous Permian mafic and ultramafic intrusions in the Kanggurtag belt, including the Huangshan, Xiangshan, and Hulu intrusions (Qin et al. [2003;](#page-20-0) Han et al. [2004;](#page-18-0) Zhou et al. [2004](#page-21-0)).

Geology of the Yamansu deposit

In the Yamansu ore district, the host rocks include the lower Carboniferous Yamansu Formation and the upper Carboniferous Shaquanzi Formation (Fig. [2](#page-3-0)). The Yamansu Formation comprises andesitic tuff and andesitic volcanic breccia intercalated with limestone in the upper unit and

volcanic and volcanic-clastic rocks in the lower unit. The upper unit hosts iron mineralization (Fig. [2\)](#page-3-0). Hou et al. ([2014b](#page-18-0)) obtained a zircon age of 324.4 ± 0.94 Ma for basalt of the Yamansu Formation in the Yamansu deposit. The Shaquanzi Formation comprises flysch intercalated with bimodal volcanic and carbonate rocks and is overlain by Lower Permian marine and terrestrial clastic rocks (Mao et al. [2005](#page-19-0)). Four major NNE- to NEE-trending faults have been recognized in the ore district (Fig. [3](#page-4-0)a). Among them, post-ore faults, F_2 and F8, crosscut and displaced the orebodies (Fig. [3](#page-4-0)).

The Yamansu deposit is intruded by some subvolcanic pyroxene-diorite porphyries, cropping out about 500 m southwest of the orebodies (Fig. [3a](#page-4-0)). There are seven N-S-trending Fig. 2 Comprehensive stratigraphic column of the Yamansu ore district (modified from Zhang and Zhang [1996](#page-21-0))

mafic dykes in the open pit of the Yamansu deposit (Li [2012\)](#page-19-0). These Permian dykes are 10 to 50 cm wide and crosscut the marble and basalt (Li [2012\)](#page-19-0). Minor pyrite and garnet skarn formed in the contact zone between the mafic dykes and marble (Li [2012\)](#page-19-0).

Skarns

Skarns are ubiquitous and intensively developed in the Yamansu deposit. A body of skarn has a strike length of \sim 1000 m, a depth in excess of 600 m, and an average width of 120 m as revealed by drilling (Fig. [3b](#page-4-0)). This skarn body has a sharp boundary with the country rocks of marble and basalt and is closely associated with the orebodies (Online Resource 1; a, b). Limestone was partly altered with obvious alteration bands nearly parallel to the sedimentary bands (Online Resource 1; c). The skarn is commonly closely associated with ores and altered volcanic rocks (Online Resource 1; d). Relics of garnet in the massive ores (Online Resource 1; e) indicate the earlier formation of garnet skarn. The garnet skarn is mainly composed of garnet with minor diopside (Online Resource 1; e), overprinted by later hydrothermal alteration consisting of epidote, chlorite, amphibole, prehnite, and pyrite (Online Resource 1; f–h).

Iron orebodies

There are more than 16 orebodies in the deposit, constituting a mineralized belt 1640 m long and 120 m wide with an area of 0.2 km^2 (Yao et al. [1993](#page-21-0)). The orebodies are hosted in calcic skarn and locally in limestones adjacent to the skarn, occurring as EW-trending bodies with stratabound or lenticular shapes (Fig. [3a](#page-4-0), b). They have abrupt contact with country rocks (Online Resource 1; a, b). No. 1, 2, 4, and 8 orebodies are the largest ones. No. 1 orebody is in the lower part of the iron ore belt and is > 940 m long, and dips southwards at 43° at surface to 72° at 560 m below (Yao et al. [1993](#page-21-0)). This orebody has a width of 1.25–19.3 m (Yao et al. [1993\)](#page-21-0). The ore grade is 20.9–63.7% total

Fig. 3 Geological (a) and cross section (b) map of the Yamansu deposit (modified from Hou et al. [2014b\)](#page-18-0)

Fe with an average grade of 47.7% (Yao et al. [1993](#page-21-0)). The No. 2 and 3 orebodies are located in the upper part of the ore belt and probably merge into one orebody, which dips southwards at 35– 55° and has a length of 1084 m, a width of 6–62 m, and an average thickness of 17.6 m. The No. $2 + 3$ orebody has a grade of 27.3–62.9% total Fe and an average grade of 54.9%. The No. 4 and 8 orebodies are located in the middle part of the ore belt. The No. 4 orebody is 108 m long and 8 m wide at the surface and dips southwards at 35–50°. The No. 8 orebody is 55 m long and 5 m wide and is 302 m at depth.

Iron-rich fragments were also identified in association with volcanic or submarine volcanic breccia and ignimbrite adjacent to the orebodies (Li et al. [2015](#page-19-0)). These fragments display typical volcanic fabric features, such as porphyritic texture, and hyalopilitic texture of the groundmass and vesicles filled by minerals to form amygdales. Iron-rich fragments are composed of mainly magnetite with various amounts of oligoclase, albite, and K-feldspar. These fragments are characterized by high Si, Al, Fe, Ca, Ti, Na, and K contents and low Mg content with the average total Fe content of 26% (Li et al.

[2015\)](#page-19-0). Magnetite in these fragments is mostly Ti-V-rich magnetite, with the TiO₂ content ranging up to 4.9 wt% and V_2O_3 content up to 3.2 wt\% (Li et al. [2015\)](#page-19-0).

Mineralization

Magnetite and pyrite are the dominant ore minerals. Chalcopyrite is not common and is commonly altered to malachite. Gangue minerals include garnet, hornblende, biotite, chlorite, epidote, quartz, calcite, and other calc-silicate minerals. Jarosite, malachite, and gypsum are supergene in origin. Ores are mainly massive and banded, but disseminated ores and ore veinlets are locally present. Massive ores commonly contain 80–90 modal% magnetite and 10–20 modal% pyrite, calcite, or garnet (Online Resource 2; a–e). Most of the massive ores are crosscut by calcite or quartz veins (Online Resource 2; a, d). Some massive ores are poorly consolidated with pores filled by pyrite and chlorite (Online Resource 2; e). Banded ores consist of 60–80 modal% magnetite and 20–40 modal% garnet and/or pyrite (Online Resource 2; f-h). Bands of magnetite range from a few millimeters to about 5 cm (Online Resource 2; f–h). The boundaries between bands are sharp (Online Resource 1; e and Online Resource 2; f) or gradational (Online Resource 2; g, h). Disseminated ores have fine-grained magnetite in skarn or wall rocks (Online Resource 2; i). Magnetite and/or hematite veins of variable width crosscut the early garnet skarn (Online Resource 2; j, k) or marble (Online Resource 2; l).

Paragenetic sequence of alteration and mineralization

Extensive alteration surrounding the orebodies includes skarn, epidote, chlorite, pyrite, carbonate, and silicate alteration from early to late. Skarn is the main alteration type closely related to iron mineralization. Based on the macro- and microtextural relationships of mineral assemblages, four principal stages of alteration and mineralization are identified in the Yamansu deposit (Fig. [4\)](#page-6-0). The paragenetic sequence includes garnetdiopside (I), magnetite (II), hydrous silicate-sulfide (III), and calcite-quartz (IV) stages.

Stage I: garnet-diopside stage

This stage is dominated by garnet and diopside, similar to the typical prograde stage of skarn deposits. In most cases, garnet is more common and diopside can only be identified under the microscope. Garnet shows dark brown, red brown, and light brown colors from early to late (Online Resource 1; d, e) and has a grain size ranging from a few millimeters to centimeters (Online Resource 2; c, i, j). The different colors of garnet represent different Fe and Al contents. Electron microprobe analyses show that garnet in the Yamansu deposit belongs to the andradite-grossular series (Li [2012](#page-19-0); Li et al. [2014](#page-19-0)). The formation temperature of garnet is 310–800 °C (Zhang and Zhang [1996](#page-21-0)). Garnet is contemporary with magnetite (Online Resource 2; c and Fig. [5a](#page-7-0)) or forms slightly earlier than magnetite (Online Resource 2; g, i, j and Fig. [5](#page-7-0)b, c). Diopside commonly occurs as radial aggregates associated with garnet (Fig. [5a](#page-7-0)). This stage is similar to the prograde stage of skarn Fe mineralization elsewhere (Meinert et al. [2005\)](#page-19-0).

Stage II: magnetite stage

This is the main stage of Fe mineralization. Magnetite is the dominant ore mineral associated with garnet and diopside (Fig. [4](#page-6-0)). Magnetite occurs as massive aggregates, disseminated grains, or veins (Online Resource 2). Magnetite typically occurs as subhedral to euhedral grains ranging from 10 to 500 μ m in length (Fig. [5a](#page-7-0)–d). Some magnetite grains are radial and rounded (Fig. [5b](#page-7-0), d) and are locally replaced by hematite (Fig. [5](#page-7-0)d). Some magnetite grains show compositional zoning and are composed of dark and light domains in backscattered electron (BSE) images (Fig. [5](#page-7-0)e). Quartz and calcite associated with magnetite have fluid inclusion homogenization temperatures of 362 to 476 °C (Zhang and Zhang [1996](#page-21-0)).

Stage III: hydrous silicate-sulfide stage

This stage is characterized by the wide occurrence of pyrite, chalcopyrite, actinolite, epidote, chlorite, and prehnite (Online Resource 2; a, b, f and Fig. [6\)](#page-8-0). Pyrite occurs as veins with varying width crosscutting massive ores (Online Resource 2; a, f) or occurs as massive aggregates associated with magnetite (Online Resource 2; b, g). Pyrite is also disseminated in massive and banded ores (Online Resource 2; d, e). Pyrite occurs as subhedral to euhedral grains ranging from 10 to 1000 μm in diameter (Fig. [5](#page-7-0)f–h). Chalcopyrite is rare and associated with pyrite, actinolite, and quartz (Fig. [5f](#page-7-0), g). Part of chalcopyrite grains has been oxidized to malachite (Fig. [5](#page-7-0)h). Epidote and chlorite commonly occur as veins in massive ores (Online Resource 2; a), indicating their later formation than magnetite. Epidote and chlorite are also widespread in altered volcanic rocks (Online Resource 1; f–h). The calculated temperature for chlorite formation is about 280 °C using the chlorite geothermometer (Li [2012\)](#page-19-0). This stage is similar to the retrograde stage of skarn Fe mineralization elsewhere (Meinert et al. [2005](#page-19-0)).

Fig. 4 Simplified paragenetic sequence of mineralization and alteration in the Yamansu deposit (modified from Li et al. [1986\)](#page-19-0)

 \bigstar Magnetite used for trace element analysis

→ Pyrite used for Re-Os dating and S isotope analysis

 \triangle Calcite used for C and O isotope analysis

Stage IV: calcite-quartz stage

Later calcite \pm quartz veins are widespread in the Yamansu deposit. Calcite \pm quartz veins commonly crosscut early ores (Online Resource 2; a, d, f), indicating their later formation. The veins have the width ranging from a few millimeters to 2 cm.

Analytical methods

Re-Os isotope analyses

Eleven pyrite separates from the magnetite-sulfide ores were selected for Re-Os isotope analysis. Ores were crushed and sieved, and magnetite was removed by a magnet. Pyrite grains were separated by heavy liquid and then hand-picked under a binocular microscope. The grains range in size from 150 to 250 μm. Re and Os isotopes were analyzed using a PE ELAN DRC-e ICP-MS in the State Key Laboratory of Ore Deposit Geochemistry, Institute of Geochemistry, Chinese Academy of Sciences (IGCAS), Guiyang, China. The detailed analytical procedure was described in Qi et al. [\(2013\)](#page-20-0). Approximately 0.05 to 0.1 g of pyrite was weighed and loaded into a 200-ml Carius tube with known amounts of ¹⁸⁵Re and ¹⁹⁰Os spikes. Samples were then digested and equilibrated using 10 ml concentrated HNO₃ and 2 ml HCl at 200 $^{\circ}$ C for about 12 h. Osmium was separated as $OsO₄$ from the matrix using in situ distillation equipment and Re was separated from the remaining solution after Os distillation using anion exchange resin (Biorad AG 1 × 8, 200–400 mesh) (Qi et al. [2007](#page-20-0); Qi et al. [2010\)](#page-20-0). Iridium was added to Re and Os solutions for mass discrimination correction (Schoenberg et al. [2000;](#page-20-0) Huang et al. [2013b\)](#page-19-0). Procedural blanks were 7.0 ± 1.0 pg $(n=2)$

Fig. 5 Photomicrographs of ores in the Yamansu deposit. a Euhedral magnetite associated with garnet, calcite, quartz, and diopside (BSE image). b Needle-like or foliaceous magnetite associated with zoned garnet (BSE image). c Subhedral to anhedral magnetite between euhedral zoned garnets indicates the later formation of magnetite. d Circular coarse-grained magnetite surrounded by anhedral fine-grained magnetite. Magnetite partly replaced by hematite. e Single magnetite grain composed of dark and light domains (BSE image). f Fine-grained magnetite associated with pyrite, chalcopyrite, actinolite, and quartz (BSE image). g Fine-grained magnetite replaced by pyrite and chalcopyrite. h Chalcopyrite partly oxidized to malachite along the grain margin. Mineral abbreviations: Act actinolite, Cal calcite, Ccp chalcopyrite, Chl chlorite, Di diopside, Ep epidote, Grt garnet, Hem hematite, Mag magnetite, Mal malachite, Py pyrite, Qz quartz \blacktriangleleft

and 1.5 ± 0.1 pg ($n = 2$) for Re and Os, respectively, with an average $187Os/188Os$ value of 0.55 ± 0.10 (1σ , $n = 2$). Absolute uncertainties (2σ) are derived from error propagation of uncertainties in Re and Os mass spectrometer measurements, blank abundances and isotopic compositions, and spike calibrations.

Stable isotope analyses

Sulfur isotopes were analyzed using a EuroVector EA3000 element analyzer (EA) coupled to a GV IsoPrime IRMS (EA-IRMS) at the State Key Laboratory of Environmental Geochemistry, IGCAS. The method was similar to that described in Huang et al. ([2013b](#page-19-0)). The sulfur isotopic compositions are reported as δ -notation in per mil (‰) relative to Vienna-Canyon Diablo Troilite (VCDT). Replicate analyses of Chinese Ag₂S standard GBW04415 yielded δ^{34} S values of $22.13 \pm 0.08\%$ (1 σ , n = 12), in good agreement with the certified values of $22.15 \pm 0.14\%$ (1 σ) (Ding et al. [2001](#page-18-0)).

Carbon and oxygen isotopic compositions were determined at the State Key Laboratory of Ore Deposit Geochemistry, IGCAS, using a GasBench II device in the continuous flow mode coupled with a MAT 253 Isotope Ratio Mass Spectrometer (IRMS) following the equilibration method (Rangarajan and Ghosh [2011\)](#page-20-0). The carbon dioxide gas was extracted by reaction with 15 drops of pure phosphoric acid in a sealed quartz tube at 72 °C for 1 h and carried to a mass spectrometer by helium gas. Carbon and oxygen isotope ratios afforded by IRMS are expressed as δ -notation in per mil (‰) relative to Vienna-Pee Dee Belemnite (VPDB). The oxygen isotope ratios are usually reported to be relative to Vienna-Standard Mean Ocean Water (VSMOW) standard. Duplicate analyses of standards have a precision better than 0.2‰.

Electron probe microanalyses

Major and selected trace elements in magnetite were measured using a JEOL JXA-8230 Electron Probe Micro-Analyzer (EPMA) at the Testing Center of Shandong Bureau of China

Fig. 6 a Plot of 187 Re/ 188 Os versus 187 Os/ 188 Os for all pyrite samples. b Plot of 187 Re versus 187 Os for LLHR pyrite. c Weighted average of Re-Os model ages for LLHR pyrite. Regression of all data was performed using Isoplot version 3.27 (Ludwig [2003](#page-19-0)). Error correlation coefficient (rho) was added during regression of isotope ratios. MSWD mean square of weighted deviate

Metallurgical Geology Bureau, Jinan, China. Elements were analyzed using an accelerating voltage of 15 kV, a beam current of 20 nA, and a beam spot of 1 μm. Matrix effects were corrected using the ZAF software provided by JEOL. The accuracy of the reported values for the analyses is 1–5% depending on the abundance of the element. The detailed operating conditions and mean detection limits of elements are listed in Online Resource 3. In addition to quantitative spot analyses, wavelength dispersive X-ray (WDX) maps were collected using an accelerating voltage of 20 kV, a probe current of 100 nA, point size of 0.3×0.3 µm, and dwell time of 5 ms for acquisition on each point.

Laser ablation ICP-MS trace element analyses

Thin sections of ore samples were examined by electron microprobe to characterize the textures of magnetite prior to LA-ICP-MS analysis. Major and trace elements of magnetite were determined by a Coherent GeoLasPro 193 nm laser ablation system coupled with an Agilent 7700x ICP-MS at the State Key Laboratory of Ore Deposit Geochemistry, IGCAS. Helium was applied as a carrier gas and argon was used as the make-up gas and mixed with the carrier gas via a T-connector before entering the ICP. Each analysis includes a background acquisition of approximately 20 s (gas blank) followed by 40 s data acquisition from the sample. Analytical spots (32 μm) were ablated by 160 successive laser pulses (4 Hz). Every 8 sample analyses were followed by one analysis of GSE-1G as quality control to correct the time-dependent drift of sensitivity and mass discrimination. A komatiitic glass (GOR-128g) and a natural magnetite (BC28) from the Bushveld Complex were also used to monitor the data quality by analyzing it as an unknown sample each run. Off-line selection and integration of background and analyte signals, and time-drift correction and quantitative calibration were performed by ICPMSDataCal (Liu et al. [2008](#page-19-0)). Element contents were calibrated against multiple reference materials (GSE-1G, BCR-2G, BIR-1G, BHVO-2G, and NIST 610) using 57 Fe as internal standard (Liu et al. [2008](#page-19-0); Gao et al. [2013\)](#page-18-0). The analytical results of monitoring samples BC28 and GOR-128g are consistent with their certified values and reported values from previous studies (Online Resource 3).

Results

Re-Os isotope composition

The Re-Os isotope composition of pyrite samples is listed in Online Resource 4. Pyrite has total Re and Os contents ranging from 105 to 426 ppb and from 0.432 to 1.193 ppb, respectively (Online Resource 4). Pyrite has small amounts of common Os contents of 0.026–0.776 ppb (Online Resource 4). As a result, they have highly variable 187 Re/ 188 Os and 187 Os/ 188 Os values of 669–43,479 and 3.8–235, respectively. Due to low common Os, the blank 188 Os accounts for a significant component of the measured value, which would result

in highly correlated uncertainties between 187 Re/ 188 Os and $187Os/188Os$ values. To correct this effect, the $187Re/188Os$ and $187Os/188Os$ data are plotted with the associated uncertain-ty correlation value (rho; 0.29–0.96) (Ludwig [1980\)](#page-19-0) and 2σ uncertainties. Conventional plot of 187 Re/ 188 Os versus $187Os/188Os$ regression analysis for all pyrite samples yields a model 1 isochron age of 321.7 ± 7.3 Ma and an initial 187 Os/ 188 Os value of 0.22 ± 0.28 (2 σ , MSWD = 0.51, Fig. [6](#page-8-0)a). The 187Os budget is predominately composed of radiogenic 187 Os (187 Os^r; 94.3–99.9%) and, therefore, defines these minerals as low-level highly radiogenic (LLHR) sulfides (Stein et al. [2000](#page-20-0)). For LLHR sulfides, a direct plot of 187 Re and 187Os contents is preferred. Regression of eight Re-Os analyses with 187 Re/ 188 Os values larger than 5000 yields a model 1 isochron age of 319 ± 22 Ma, with the initial ¹⁸⁷Os content of 0.018 ± 0.042 ppb $(2\sigma, MSWD = 0.25; Fig. 6b)$ $(2\sigma, MSWD = 0.25; Fig. 6b)$ $(2\sigma, MSWD = 0.25; Fig. 6b)$. For LLHR sulfides, model ages can be calculated using 187 Re and 187 Os^r contents. Re-Os model ages of these pyrite samples range from 319 to 337 Ma (Online Resource 4), and the weighted average of these model ages is 326.7 ± 7.3 Ma (2σ , MSWD = 0.30, probability = 0.95; Fig. [6](#page-8-0)c), consistent with the isochron age within errors.

Sulfur isotope composition

Pyrite has $\delta^{34}S_{\text{VCDT}}$ values ranging from -2.2 to +2.9‰, but mostly between − 1 and 0‰ (Online Resource 4 and Fig. 7). The sulfur isotope composition of H_2S in equilibrium with pyrite can be calculated from the $\delta^{34}S_{\text{VCDT}}$ values of pyrite, appropriate fractionation factors between aqueous sulfur species and sulfide minerals, which are closely related to temperature and $fO₂$ during sulfide precipitation (Ohmoto and Rye [1979\)](#page-20-0). Fluid inclusion homogenization temperatures of calcite associated with pyrite from this stage range from 240 to 380 °C (Zhang and

Fig. 7 Histograms showing sulfur isotope composition of pyrite from the Yamansu deposit and H_2S in fluids in equilibrium with pyrite

Zhang [1996\)](#page-21-0). The sulfur isotope composition of pyrite can represent the sulfur composition of total sulfur in the hydrothermal fluids because of lack of coexisting sulfates (Kajiwara [1971](#page-19-0)). The calculated sulfur isotope composition of H₂S has δ^{34} S_{VCDT} values ranging from − 3.1 to + 2‰ with an average of -0.8% (Online Resource 4).

Carbon and oxygen isotope composition

Calcite from stage II has $\delta^{13}C_{\text{VPDB}}$ and $\delta^{18}O_{\text{VSMOW}}$ values of – 2.5 to -1.2% and 11.8 to 12.0‰, respectively, whereas those from stage IV have values of -1.1 to 1.1% and -7.7 to -5.2% . respectively (Online Resource 4). The $CO₂$ in equilibrium with stage II calcite has C and O isotope composition of − 1.1 to 0.2‰ and 11.7 to 11.9‰, respectively, whereas $CO₂$ in equilibrium with stage IV calcite has these values of -4.4 to $-2.2%$ and -22.2 to -19.7% , respectively. In the plot of δ^{13} C_{VPDB} versus δ^{18} O_{VSMOW}, calcite from stage II lies in the field of primary magmatic carbonate, similar to other Fe deposits in the same belt (Fig. 8a). Carbon in stages II and IV has an isotope composition

close to mantle $CO₂$ (Fig. 8b). Stage II fluids have O isotope composition slightly higher than magmatic water, close to metamorphic and formation water compositions, whereas stage IV fluids have O isotope composition typical of admixture of meteoric water (Fig. 8c).

Chemical zoning of magnetite

In the Yamansu deposit, some magnetite grains have two main domains under BSE imaging (Figs. [5e](#page-7-0) and [9](#page-11-0)). One domain is dark and smooth, whereas the other domain is light and sometimes porous (Fig. [9\)](#page-11-0). Light and dark domains in the same magnetite grain can form obviously oscillatory zoning or sector zoning (Fig. [9](#page-11-0)a, b). Dark domains can be replaced to various extent by light domains along fractures and/or grain boundaries that have appreciable porosities (Fig. [9](#page-11-0)c, d), whereas light domains are generally in sharp contact with dark domains and contain numerous solid inclusions such as carbonate and silicate minerals (Fig. [9](#page-11-0)c). Single

Fig. 8 a $\delta^{13}C_{\rm PDB}\delta^{18}O_{\rm SMOW}$ diagram of calcite from the Yamansu Fe deposit. Data sources: mantle-derived rocks, granitic rocks, primary carbonates, marine carbonates, and sedimentary organic carbon (Ohmoto [1986](#page-20-0); Valley [1986;](#page-21-0) Nelson et al. [1988](#page-20-0); Hoefs [1997](#page-18-0); Ray et al. [1999](#page-20-0)). **b**, **c** δ^{13} C_{PDB} (after Clark and Fritz [1997](#page-18-0)) and $\delta^{18}O_{SMOW}$ (after Hoefs [1997](#page-18-0)) values of calcite and representative geological reservoirs. Calcite from the Heifengshan Fe deposit and Shaquanzi Fe-Cu deposit in the same belt with the Yamansu deposit are also shown for comparison (Huang et al. [2013a\)](#page-19-0)

Fig. 9 BSE images showing microtextures of magnetite. a Euhedral zoned magnetite composed of dark and light domains. The dark domains show obvious oscillatory zoning. b Subhedral magnetite grains composed of light and dark domains, forming zoned magnetite with sector zoning. c Subhedral to anhedral magnetite grains composed of early dark and late light domains. Light domains are porous. d Subhedral magnetite grains composed of light domain in the rim and dark domain in the center. e, f Euhedral magnetite grain composed of light and dark domains

magnetite grains may show obvious compositional zoning of dark and light domains (Fig. 9e, f). EPMA mapping shows that light magnetite has higher Fe content but lower Al, Si, and Mg contents than dark magnetite (Fig. [10\)](#page-12-0). Quantitative analyses of 46 spots indicate that light magnetite has higher Cr but lower Mn, Ca, and Ti contents than dark magnetite (Online Resource 4 and Fig. [11](#page-13-0)a). Due to different Ca+Al+Mn and Ti+V contents, light and dark magnetite domains define different populations (Fig. [11b](#page-13-0)).

LA-ICP-MS trace element compositions of magnetite

Due to the small scale zoning, light and dark domains cannot be separated during laser ablation, and thus, the LA-ICP-MS results represent the mixed composition of light and dark domains. In this study, only magnetite grains without zoned textures are analyzed. One

hundred and twenty-nine analyses were obtained. The average values and standard deviations of trace element contents of magnetite are listed in Online Resource 4. Full analytical results of LA-ICP-MS for different groups of magnetite are presented in Online Resource 5. Forty-seven elements were analyzed but only 28 elements were detected. Elements, such as Sc, P, Ge, Nb, Hf, Ta, Ag, Cd, In, and Au, are below the detection limits. Some magnetite grains also have Cr, Ni, Cu, As, Rb, Y, Zr, Mo, Sn, Sb, Ba, W, Bi, and Th contents lower than the detection limits.

Most element contents vary within one or two orders of magnitude, but some elements such as V, Y, Mo, W, Pb, and U vary up to three or four orders of magnitude (Fig. [12](#page-13-0)). Most magnetite grains have Mg, Al, Si, Ca, Ti, and Mn contents between \sim 100 and 10,000 ppm; V, Cr, Co, Ni, Zn, Ga, and As contents between \sim 1 and 100 ppm; and Y, Zr, Mo, Sn, Sb, Bi, Th, and U contents lower than

Fig. 10 a, b BSE images of magnetite grains and corresponding wavelength dispersive X-ray maps of selected trace elements in magnetite from the Yamansu deposit. The dark and light domains of magnetite grains have different Fe, Al, Si, and Mg contents

10 ppm (Fig. [12](#page-13-0)). Several elements show good correlation. Aluminum shows positive correlations with Mg, Mn, and Ga (Fig. [13a](#page-14-0)–c). There are also positive correlations of Si with Ca and Mg (Fig. [13d](#page-14-0), e). Antimony shows positive correlations with As and Bi (Fig. [13f](#page-14-0), g). Tungsten shows a positive correlation with Mo (Fig. [13h](#page-14-0)).

Fig. 11 a Binary plots showing EPMA data of dark magnetite and light magnetite. Dark magnetite has lower Fe and Cr contents but higher Al, Si, Mg, Mn, Ti, and Ca contents than light magnetite. b Plot of Ca+Al+Mn versus Ti+V showing skarn-like dark magnetite. c Plot of FeO versus SiO2 showing varying compositions between dark magnetite and light magnetite. Data points are individual analyses. Reference fields are from Dupuis and Beaudoin [\(2011\)](#page-18-0). BIF banded iron formation, Skarn Fe-Cu skarn deposits, IOCG iron oxide-copper-gold deposits, Porphyry porphyry Cu deposits, Kiruna Kiruna apatite-magnetite deposits, Fe-Ti, V magmatic Fe-Ti-oxide deposits

Fig. 12 Box and whisker plot for LA-ICP-MS data of magnetite grains without zoned textures. Boxes outline the 25th to 75th percentiles and whiskers extend to the minimum and maximum values. The whisker represents the range of trace element contents. The short line in the box represents the median value, whereas the circle filled white means the average value. Trace element contents below the limit of detection were removed from the box and whisker plot

Discussion

Timing of iron mineralization

Li and Chen [\(2003](#page-19-0)) reported a Rb-Sr isochron age of $374 \pm$ 44 Ma for andesitic porphyry and a garnet-epidote Sm-Nd isochron age of 352 ± 46 Ma, which were thought to represent the timing of volcanism and iron mineralization, respectively. However, these ages have large errors. Hou et al. [\(2014b](#page-18-0)) reported a LA-ICP-MS U-Pb age of 323.5 ± 1.0 Ma for hydrothermal zircon in the skarn, which was interpreted as the age of iron mineralization.

Magnetite is the main ore mineral in Fe deposits, and thus, Re-Os dating of magnetite can provide a direct mineralization age (e.g., Mathur et al. [2002;](#page-19-0) Davies [2010;](#page-18-0) Huang et al. [2014b](#page-19-0); Barra et al. [2017](#page-18-0)). However, magnetite in Yamansu ores has Re contents commonly lower than 1 ppb (unpublished data), which make it difficult to date magnetite using the Re-Os isotope method. Alternatively, the Re-Os isotope system of pyrite has been proved to be a reliable tool to date magmatic-hydrothermal events due to the relatively high Re in pyrite and the high closure temperature of this isotope system (Brenan et al. [2000](#page-18-0); Morelli et al. [2004](#page-20-0); Selby et al. [2009;](#page-20-0) Huang et al. [2013a](#page-19-0), [c](#page-19-0)). In order to obtain geologically meaningful ages, the occurrence and the paragenetic sequence of pyrite should be clearly understood. Pyrite for our Re-Os isotope analyses is closely associated with magnetite and formed simultaneously with magnetite or slightly after according to ore textures (Online Resource 2). Moreover, previous studies showed that Re-Os dating of pyrite and associated magnetite from the Shaquanzi magmatic-hydrothermal Fe-Cu deposit in the ETOB yielded similar age and demonstrated that the age of pyrite can represent the timing of iron

Fig. 13 Scattergrams for a selection of trace elements in magnetite without zoned textures. a–c Aluminum correlates positively with Mg, Mn, and Ga. d, e Silicon shows positive correlations with Mg and Ca. f, g Antimony shows positive correlations with As and Bi. h Tungsten shows a positive correlation with Mo

mineralization (Huang et al. [2014b\)](#page-19-0). Pyrite separates from the Yamansu ores have a Re-Os isochron age of 321.7 ± 7.3 Ma (Fig. [6a](#page-8-0)), which represents the age of magnetite formation. Therefore, we suggest that the Yamansu deposit formed in the Early Carboniferous.

Formation of chemical zoning in magnetite

In the Yamansu deposit, the chemical zoning of magnetite is commonly composed of two types of magnetite. The dark magnetite is inclusion-free and has darker BSE contrast,

whereas the light magnetite is sometimes porous and has lighter BSE contrast (Fig. [9\)](#page-11-0). Some magnetite grains exhibit obvious oscillatory zoning composed of light and dark bands of varying width (Fig. [9a](#page-11-0), b), similar to magnetite from the Ocna de Fier-Dognecea Fe-Cu-Pb-Zn skarn deposit in Romania (Ciobanu and Cook [2004](#page-18-0)), the Fe skarn of Vegas Peledas in Argentina (Dare et al. [2014](#page-18-0)), the lava flows of EI Laco in Chile (Dare et al. [2015\)](#page-18-0), and the brecciated diorite intrusion in the Los Colorados iron oxide-apatite deposit of Chile (Knipping et al. [2015b\)](#page-19-0). The oscillatory zoning in magnetite is commonly interpreted as growth zoning forming as a result of changes in fluid compositions and/or physicochemical parameters (such as temperature and oxygen fugacity) during crystal growth, which periodically affect the partitioning behavior of trace elements into magnetite and co-forming minerals (Shimazaki [1998;](#page-20-0) Ciobanu and Cook [2004](#page-18-0); Dare et al. [2015;](#page-18-0) Knipping et al. [2015b](#page-19-0); Makvandi et al. [2015](#page-19-0)). Therefore, oscillatory zoning in zoned magnetite from the Yamansu deposit was likely derived from a fluctuating fluid composition and/or physicochemical conditions.

In addition to zoned magnetite, some magnetite also shows irregular compositional domains that include early dark Si-rich domain and late light Si-poor domain (Fig. [9](#page-11-0)c–f), which are very common in Fe skarn deposits (Ciobanu and Cook [2004;](#page-18-0) Hu et al. [2014,](#page-19-0) [2015;](#page-19-0) Zheng et al. [2017\)](#page-21-0), iron oxide-apatite (IOA) deposits (Nold et al. [2014](#page-20-0); Knipping et al. [2015a](#page-19-0), [b\)](#page-19-0), and volcanogenic massive sulfide (VMS) deposits (Westendorp et al. [1991](#page-21-0); Makvandi et al. [2016\)](#page-19-0). Si-rich domains of magnetite grains have higher Mg, Al, Ca, Mn, and Ti contents, but lower Fe and Cr contents than Si-poor domains (Figs. [10](#page-12-0) and [11](#page-13-0)). Knipping et al. [\(2015a,](#page-19-0) [b\)](#page-19-0) showed that magnetite from the Los Colorados IOA deposit (Chile) has a Ti-rich core, magmatic in origin, and a Tipoor rim formed as a result of hydrothermal alteration of magmatic magnetite. Makvandi et al. [\(2016\)](#page-19-0) interpreted the Si-free core of magnetite from the West Ansil bimodal-mafic VMS deposit in Quebec (Canada), as likely formed during replacement of massive sulfides by magnetite, whereas the Si-rich rim could be either the consequence of the replacement of pre-existing quartz in the massive pyrrhotite ore, or the signature of Si-rich fluids. Zheng et al. [\(2017\)](#page-21-0) considered that the compositional variations of zoned magnetite (Ca-Si-rich rim and Ti-rich core) from the Cihai Fe skarn deposit in Xinjiang (China) resulted from the changes of fluid temperature and salinity that affected the solubility of Fe and Ti.

Dissolution-precipitation processes are commonly proposed to explain the chemical variations of different generations/types of magnetite (Putnis [2009](#page-20-0); Putnis and John [2010;](#page-20-0) Dare et al. [2015;](#page-18-0) Hu et al. [2015](#page-19-0); Heidarian et al. [2016](#page-18-0)). The dark magnetite has relatively higher contents of Si, Al, Ca, and Mg, which might be related to a major contribution of magmatic-derived fluids due to a large fluid-rock ratio at the site of mineral deposition (Nadoll et al. [2012](#page-20-0), [2014](#page-20-0); Dare et al. [2014](#page-18-0)). The dark magnetite was then extensively replaced by light magnetite preferentially along

microfractures within or along grain margins (Fig. [9](#page-11-0)c, d). The replacement reactions were facilitated by a fluid-assisted process with dissolution of early magnetite and precipitation of a secondary variety near the reaction front, with Si, Al, Ca, Mg, Mn, and Ti being significantly removed from the early magnetite during the reactions (Hu et al. [2015](#page-19-0)).

Dissolution and precipitation reactions can be affected by fluid composition, solubility of mineral phases, temperature and pressure, Eh, and pH (Whitney et al. [1985;](#page-21-0) Hemley and Hunt [1992;](#page-18-0) Putnis [2002;](#page-20-0) Putnis and Putnis [2007](#page-20-0)). It is suggested that the higher contents of small ionic radius cations such as $Si⁴⁺$ in early magnetite make the magnetite more susceptible to fluidassisted alteration because incorporation of these elements in magnetite would cause lattice defects or deformation (Huberty et al. [2012](#page-19-0); Hu et al. [2014](#page-19-0)). The isomorphous replacement of $Fe³⁺$ by $Si⁴⁺$ in the tetrahedral sites of Si-rich magnetite in the Yamansu deposit has been demonstrated in recent studies (Li et al. [2017](#page-19-0)). Infiltration of mixed basinal brine and meteoric water that had dissolved evaporites can be an important process changing the physicochemical parameters of hydrothermal fluids, which led to the dissolution and precipitation of magnetite (Hu et al. [2014](#page-19-0); Heidarian et al. [2016\)](#page-18-0). These high-salinity, Cl[−] rich ore-forming fluids enhance Fe solubility and disequilibrium between magnetite and evolving surficial fluids (Chou and Eugster [1977;](#page-18-0) Whitney et al. [1985;](#page-21-0) Hemley and Hunt [1992\)](#page-18-0). However, pyrite associated with magnetite has δ^{34} S values ranging from -2.2 to $+2.9\%$, and H₂S in equilibrium with pyrite has δ^{34} S values of − 3.1 to + 2‰, precluding dissolved evaporite in the hydrothermal fluids. Alternatively, the variations in physicochemical conditions of fluids may be more important, which has been demonstrated by oscillatory zoning in magnetite (Fig. [9a](#page-11-0), b). Carbon and oxygen isotope compositions of calcite suggest the involvement of meteoric water in the late mineralization process. Therefore, further dilution of magmatic-hydrothermal fluids by meteoric water in the late mineralization stage may induce the decline of temperature and pressure, which contributes to the dissolution of early magnetite and precipitation of late magnetite.

Origin and evolution of fluids

Sulfur isotope data of pyrite and the carbon and oxygen isotope data of calcite provide important constraints on the origin and evolution of the hydrothermal fluids. Sulfur in equilibrium with pyrite has δ^{34} S values of −3.1 to +2‰ with an average of − 0.8‰, suggesting magma-derived sulfur (Ohmoto [1972](#page-20-0)). Stage II calcite has C and O isotope composition similar to primary carbonate (Fig. [8](#page-10-0)a), whereas carbon has an isotope composition close to mantle $CO₂$ (Fig. [8b](#page-10-0)), indicating the derivation of C and O from magmatic-hydrothermal fluids. Fluids in equilibrium with stage II calcite have O isotope composition slightly higher than magmatic water, similar to those of formation and metamorphic water (Fig. [8c](#page-10-0)). Fluids in equilibrium with magnetite and calcite have $\delta^{18}O_{\text{fluid}}$ values of 12.2–12.5 and 8.3–12.2‰,

respectively (Zhang and Zhang [1996](#page-21-0)), also slightly higher than those of magmatic water (6–10‰; Taylor [1986\)](#page-20-0) but similar to those of metamorphic water (5–25‰; Taylor [1974](#page-20-0)). Due to lack of regional metamorphism in the Aqishan-Yamansu belt, we infer that metamorphic water-like O isotope composition is more likely due to the fluid-rock interaction between magma-derived hydrothermal fluids with limestone because sedimentary materials are enriched in the heavy O isotope (Fig. [8a](#page-10-0)). Stage IV fluids have C isotope composition similar to average inorganic crustal $CO₂$ (Fig. [8](#page-10-0)c) but have O values similar to meteoric water (Fig. [8c](#page-10-0)), indicating that meteoric water was involved in the hydrothermal fluids during the later stage.

Pyrite samples have variable Re and common Os contents ranging from 105 to 426 ppb and from 26 to 776 ppt, respectively. Variable Re and Os contents of pyrite cannot be due to the derivation of Re and Os from different sources in view of the well-fitted isochron. This variability may reflect fluid evolution due to the change of redox conditions and ligand chemistry with time or in different portions of the deposit (e.g., Morelli et al. [2004\)](#page-20-0), or it may be due to co-precipitated sulfides such as chalcopyrite. The initial Os isotope ratio is an important parameter to evaluate the relative contributions of crust and mantle in the formation of ore deposits. The elevated $187Os/188Os$ ratio of the crust (\sim 1.0–1.3) (Walker et al. [1991](#page-21-0)), compared to the mantle (\sim 0.11–0.15) (Meisel et al. [2001\)](#page-20-0), can be used to discern between these two reservoirs. The initial $\frac{187}{\text{Os}}$ / $\frac{188}{\text{Os}}$ ratio from the pyrite isochron is 0.22 ± 0.28 (Fig. [6a](#page-8-0)). The large uncertainty of the initial ratio makes it difficult to discuss the source of Os and, by inference, ore-forming fluids. But the highest initial Os ratio is lower than 0.5 if we consider the error, which is lower than values of the typical crust. Therefore, we can infer that Os of pyrite is mainly mantle-derived and minor crustal Os may have been involved. The trace crustal Os may have derived from magma rather than sedimentary rocks by hydrothermal replacement because the Re-Os isotope system of pyrite was closed after mineralization as indicated by the isochron. This is consistent with the dominant magmatic-hydrothermal origin of oreforming fluids indicated by C, O, and S isotopes. Moreover, the mixing of magmatic-hydrothermal fluids with meteoric water in the late hydrothermal stage did not produce a mixing trend in the Re-Os data which has been documented at the Grasberg porphyry stockwork Cu-Au mineralization in Indonesia, for instance (Mathur et al. [2005\)](#page-19-0). This indicates that the dilution of magmatic-hydrothermal fluids by meteoric water did not reset but only slightly disturbed the pyrite Re-Os isotope system in terms of the large uncertainty of the initial Os isotope ratio.

The trace element composition of magnetite also records the nature of the ore-forming fluids. EPMA results show that dark magnetite has Ca+Al+Mn and Ti+V contents similar to those of skarn deposits, indicating their hydrothermal origin (Fig. [11](#page-13-0)b). Light magnetite has lower Ca+Al+Mn contents than dark magnetite and plots in the BIF field, inconsistent with the skarn origin (Fig. [11b](#page-13-0)). This indicates that the trace element composition of light magnetite is different from that of dark magnetite and thus cannot be used to discriminate the origin of magnetite. In the plots of Ca+Al+Mn versus Ti+V and Ni/(Cr+Mn) versus Ti+V (Fig. [14](#page-17-0)), magnetite grains without zoned textures lie in the skarn field, indicating a skarn origin (Fig. [14](#page-17-0)).

Iron enrichment mechanism in the Yamansu deposit

Iron accumulation in volcanic-hosted Fe deposits commonly involves the model of magmatic versus hydrothermal, or the combination of both. For instance, Knipping et al. [\(2015a,](#page-19-0) [2015b](#page-19-0)) postulated that a separated high-salinity fluid scavenged Fe in the form of $FeCl₂$ complexes from intermediate melts to form the Kiruna-type deposit of Los Colorados (Chile). Günther et al. [\(2017](#page-18-0)) proposed different origins of two groups of magnetite from the Zhibo and Chagangnuoer Fe deposits in Western Tianshan. An ortho-magmatic fluid $(\sim 800 \degree C)$ separated from an intermediate melt accounted for the early magnetite mineralization, whereas partial remobilization of iron from the proximal, ortho-magmatic orebodies and a subsequent distal reprecipitation was responsible for the later magnetite formation. These processes were driven by late-stage hydrothermal fluids originating from deep-seated granitic/granodioritic intrusions in the immediate vicinity. Both examples show that ortho-magmatic melts and fluids are important for the Fe accumulation in a volcanic-hosted arc setting.

The Yamansu deposit shares many similarities with calcic skarn Fe deposits in tectonic setting, metals, host rocks, alteration, mineral assemblages, and origin of ore-forming fluids (Online Resource 4). Skarn is widely distributed in this deposit and has a thickness of up to 50 m. The skarn displays zones with garnet in the center and chlorite-epidote-sulfide facies at the margin (Lu et al. [1995](#page-19-0)), similar to typical skarn deposits (Meinert et al. [2005\)](#page-19-0). The Yamansu deposit has an andradite-grossular assemblage and Si-rich magnetite (Li [2012](#page-19-0)), also typical of skarn deposits (Meinert et al. [2005](#page-19-0)). The ore-forming fluids are mainly magmatic water with minor meteoric water. The trace element composition of magnetite also indicates a skarn origin (Fig. [14\)](#page-17-0). Because the majority of orebodies are hosted in skarn, iron enrichment in the Yamansu deposit is similar to that in skarn deposits elsewhere.

It is noteworthy that in addition to the skarn-related orebodies, there are also iron-rich fragments in the ignimbrite, volcanic breccia, and sedimentary volcanic breccia near the orebodies (Li et al. [2015\)](#page-19-0). Iron-rich fragments are composed of mainly magnetite with various amounts of oligoclase, albite, and Kfeldspar and have an average total Fe content of 26 wt%. These fragments were probably derived from iron-rich melts produced by the Fenner-type magma evolution that iron enrichment trends in residual basaltic liquids during magma differentiation (Fenner [1929](#page-18-0); Li et al. [2015\)](#page-19-0). Magnetite in these fragments is magmatic in origin as deduced from high Ti+V contents. Li et al. ([2015](#page-19-0)) proposed that iron enrichment in Yamansu resulted from dissolution-precipitation of early magmatic magnetite by

Fig. 14 Plots of Ca+Al+Mn versus Ti+V (a) and Ni/(Cr+Mn) versus Ti+ V (b) for LA-ICP-MS data of magnetite without zoned textures. Both individual analyses and average of a single sample are shown for

late hydrothermal fluids. However, there is no evidence for the magmatic precursor in magnetite. Therefore, the iron enrichment model in the Zhibo and Chagangnuoer iron deposits cannot be used in Yamansu. By comparing the trace element composition of dark Si-rich magnetite and light Si-poor magnetite, we propose a possible model for the iron enrichment in the Yamansu deposit. Si-rich dark magnetite was possibly formed from $Si \pm Fe$ -rich magmatic-derived fluids due to large fluid-rock ratios at the temperature of skarn formation (~ 400 °C) (Meinert et al. [2005\)](#page-19-0). Infiltration or dilution of magmatic-dominant fluids by meteoric water leads to variations in temperature and pressure of fluids, which induces the dissolution of early dark magnetite and precipitation of late magnetite. At the same time, fluctuating fluid composition and physicochemical conditions also resulted in the formation of zoned magnetite with oscillatory zoning. During this process, Si, Ca, Al, and Mn were removed from the dark magnetite, and thus, FeO contents were elevated (Fig. [11](#page-13-0)b, c). The iron enrichment mechanism by dissolution and precipitation processes was also proposed for the Chenchao skarn deposit in the middle-lower Yangtze River metallogenic belt (Hu et al. [2014\)](#page-19-0). Therefore, we suggest that iron in the Yamansu deposit was mainly derived from magmatic-hydrothermal fluids similar to skarn deposits and further iron enrichment was due to dissolution and precipitation of early Si-rich skarn magnetite during the retrograde evolution of the skarn system.

Implication for regional mineralization

Previous studies concluded that hydrothermal iron mineralization in the ETOB mainly occurred in the late Carboniferous (~300 Ma) based on the ages of the Heifengshan, Shuangfengshan, and Shaquanzi Fe deposits (Huang et al. [2013a,](#page-19-0) [2014b\)](#page-19-0). The \sim 300-Ma hydrothermal event is also indicated by other iron deposits such as the Bailingshan and Hongyuntan deposits. The Bailingshan deposit has a pyrite Rb-Sr age of 293 ± 10 Ma (Wang et al. [2015](#page-21-0)), whereas the

comparison. Reference fields are from Dupuis and Beaudoin [\(2011](#page-18-0)) and the abbreviations of deposit types are similar to those in Fig. [11b](#page-13-0)

Hongyuntan deposit has a mineralization age between \sim 324 and 300 Ma constrained by quartz keratophyre wall rock and granitic intrusions (Zheng [2015\)](#page-21-0). Mao et al. [\(2005\)](#page-19-0) suggested that the skarn Fe mineralization in the ETOB may have begun in the early Carboniferous $(\sim 350 \text{ Ma})$ and ended in the late Carboniferous \sim 290 Ma). Our Re-Os dating of pyrite from the Yamansu deposit indicates a \sim 320-Ma hydrothermal event for iron mineralization. Therefore, we suggest that hydrothermal iron mineralization in the ETOB occurred in multiple episodes peaking at ca. 320 Ma and ca. 300 Ma. Magmatic activities in the ETOB are characterized by minor Permian mafic and ultramafic intrusions in the Kanggurtag belt and numerous granitic intrusions in all belts of the ETOB. These granitoids have ages ranging from \sim 230 to \sim 386 Ma (Tong et al. [2010](#page-20-0); Zhou et al. [2010\)](#page-21-0) and two peak ages of 340–310 and 300–270 Ma (Online Resource 6). The two ages of iron mineralization $(\sim$ 320 and \sim 300 Ma) overlap with the two ages of granitic magmatism, indicating that iron mineralization in ETOB is the product of intense granitic magmatic activity although skarn Fe deposits are commonly associated with gabbroic to syenitic intrusions (Online Resource 4).

Calcic Fe skarns are associated with Fe-rich plutons intruded into limestone and volcanic wall rocks (Meinert et al. [2005\)](#page-19-0). Skarns are commonly developed in the carbonated layers of the country rock and about 85% of skarns occur within 800 m of an intrusive contact. However, skarns can spread quite far from the intrusion (up to several kilometers) if there is fluid flow (Meinert [1992\)](#page-19-0). As discussed above, magnetite formation in the Yamansu deposit involves fluid infiltration processes. Skarns in Yamansu, thus, are possibly disconnected distal skarns in the pluton-related skarn system. A subsequent distal reprecipitation of hydrothermal Fe orebodies and related skarn possibly explains volcanichosted skarn deposits lacking proximal intrusions. This metallogenetic model is very useful for the understanding of similar deposits. For example, the Cihai and Cinan Fe deposits in the Beishan area adjacent to the ETOB are considered to be

magmatic (Wang et al. [2008](#page-21-0); Tang et al. [2017](#page-20-0)) or hydrothermal (skarn) in origin (Huang et al. [2013c;](#page-19-0) Zheng et al. [2017\)](#page-21-0). These deposits may have involved both magmatic and hydrothermal processes, in which hydrothermal magnetite may represent a distal Fe component.

Conclusions

The Yamansu deposit has mineral associations, related alteration, and styles of mineralization similar to Fe skarn deposits elsewhere. The deposit was formed at \sim 320 Ma and derived from dominantly magmatic-hydrothermal fluids with minor meteoric water. Some magnetite grains are composed of dark and light domains in BSE images. Light magnetite may have formed via a process of replacement of dark magnetite due to a dissolution-reprecipitation process. Dilution of magmatic-hydrothermal fluids by meteoric water could be responsible for this process. Trace element compositions of magnetite imply a skarn origin. We suggest that iron in the Yamansu deposit was mainly derived from magmatichydrothermal fluids similar to skarn deposits and further iron enrichment was due to dissolution and precipitation of early skarn magnetite by external fluids. The Yamansu deposit may represent a distal skarn deposit, in which skarn and related Fe orebodies possibly represent the distal expression of a pluton-related skarn system.

Acknowledgments Field work was helped by staff of the Yamansu deposit. Zengsheng Li and Zhihui Dai are thanked for EPMA and LA-ICP-MS analyses, respectively. Ning An and Jing Gu are thanked for stable isotope analyses. We also thank the two editors (Bernd Lehmann and Ruizhong Hu), Ryan Mathur, and an anonymous reviewer who significantly improved the manuscript.

Funding information This work was jointly supported by grants from the National Natural Science Foundation of China (41503039 and 41673050), the Chinese 973 project (2012CB416804), and the "CAS Hundred Talents" Project to J.F. Gao (Y5CJ038000).

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